

National Aeronautics and Space Administration



Fermi
Gamma-ray Space Telescope

www.nasa.gov/fermi



The Fermi LAT First Source Catalog & Using the Catalog for Analysis

Elizabeth Ferrara
FSSC

Catalog slides are modified from Jean Ballet's presentation at the 2009 Fermi Symposium

The First LAT Catalog (1FGL)

Features:

- 11 months of data 100 MeV to 100 GeV, 23.3 Ms livetime
- 10.6 M events over the whole sky (Pass 6 v3 Diffuse class)
- Improved diffuse model and calibration with respect to 0FGL
- Detection based on integrated data (not on flares)
- Precise localization

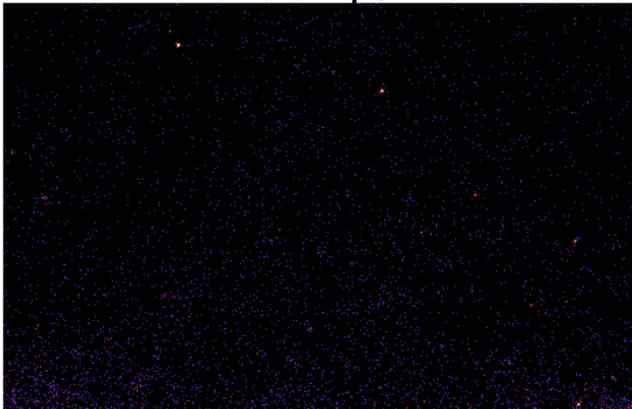
Contents:

- Source coordinates and error ellipse at 95% confidence
- Source significance and overall spectral index
- Average flux in 5 energy bands 0.1 – 0.3 – 1 – 3 – 10 – 100 GeV, plus spectral index and total flux
- Flux per month, variability index
- Quality flags: sensitivity to diffuse model, low source to bkg ratio, confusion, error ellipse not well defined, etc.
- Associations with known sources

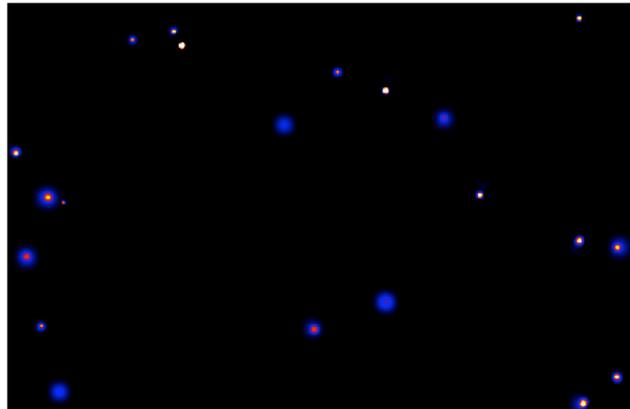
Source detection

- **Difficulty is that point spread function improves enormously from 5° at 100 MeV to nearly 0.1° above 10 GeV.**
- **Combine several energy bands, merge seeds from several detection methods.**

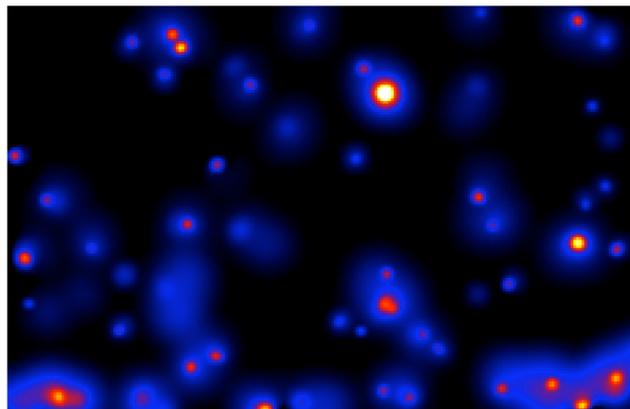
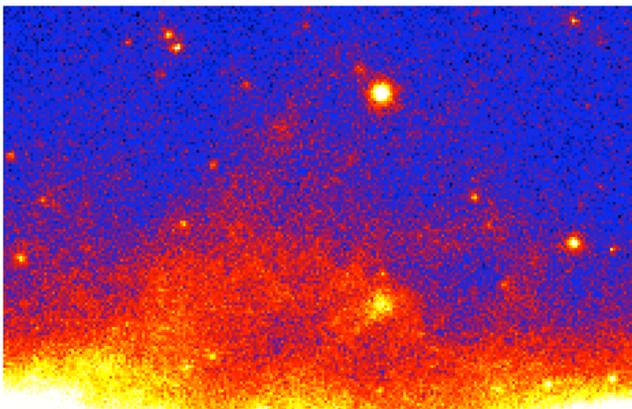
Raw counts map, $75 \times 45^\circ$



Wavelet filtered



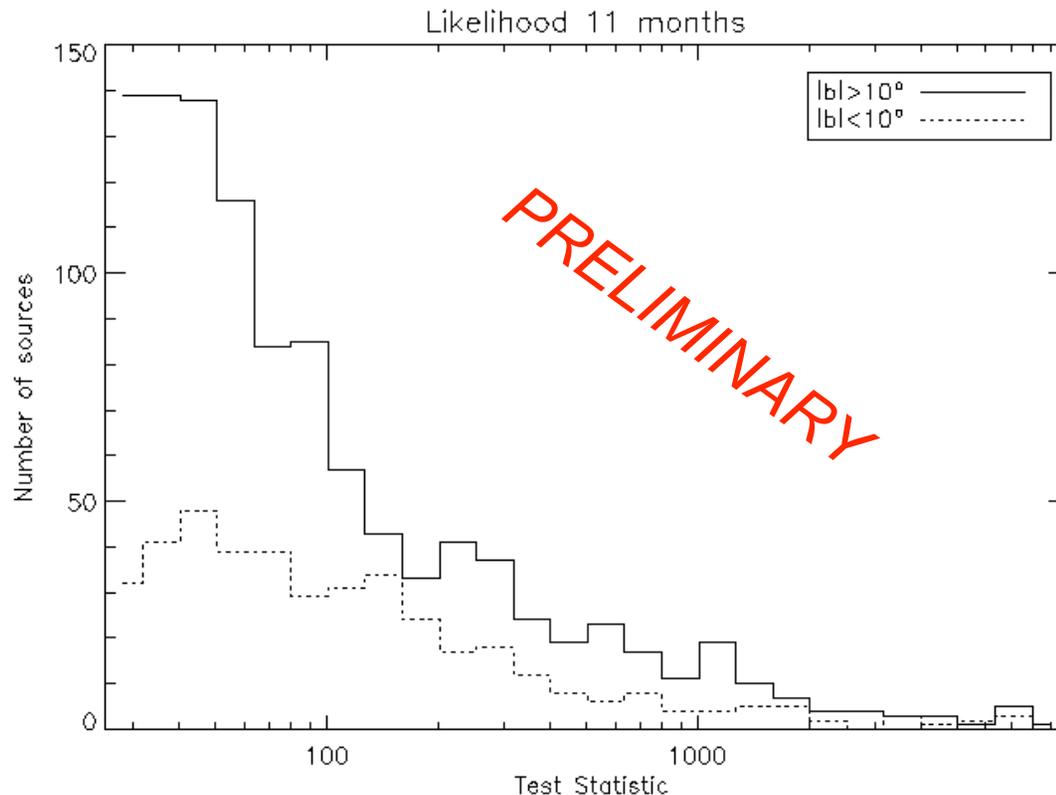
Front > 5 GeV
Back > 10 GeV
Very few events,
very well localized



Front > 200 MeV
Back > 400 MeV
Many events, not so
well localized

Source significance

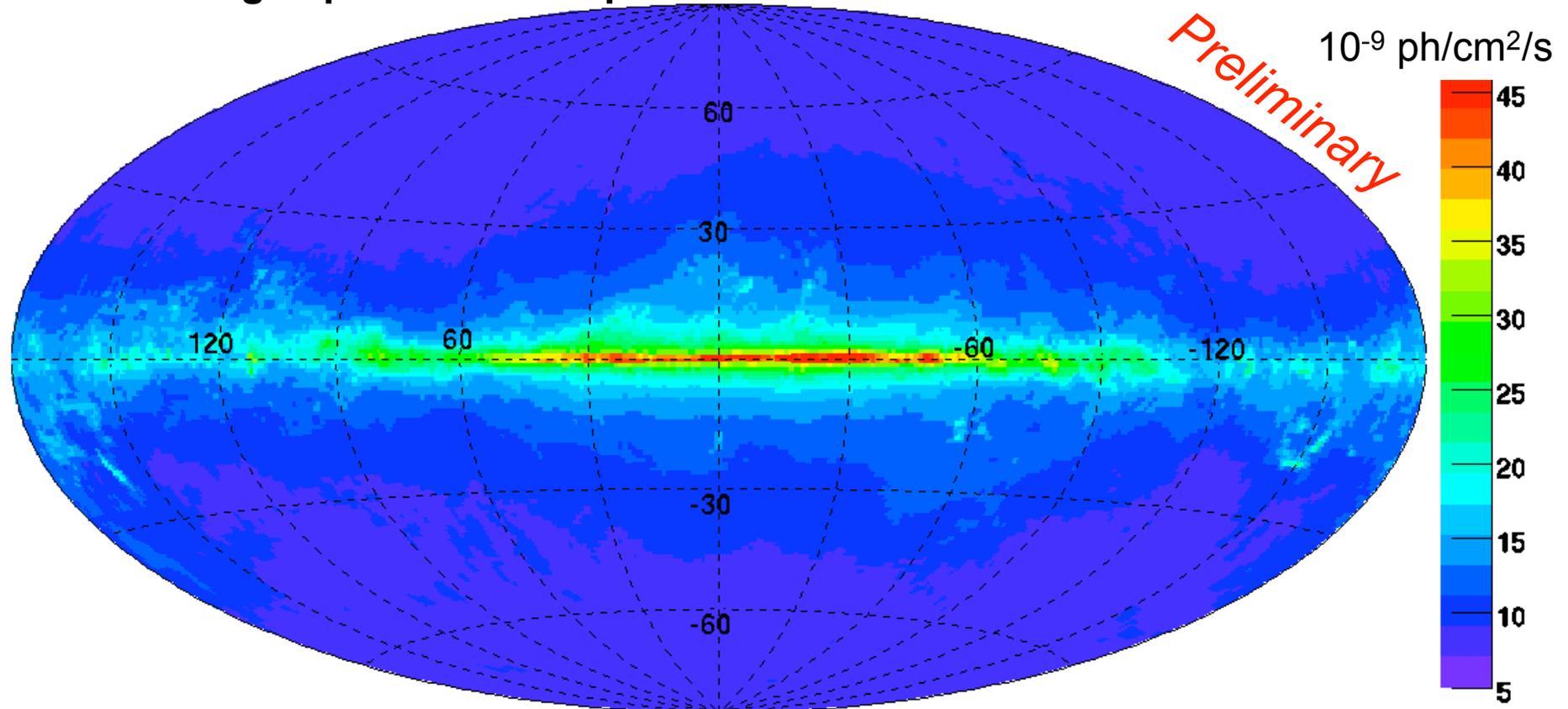
- **3D maximum likelihood analysis (position and energy) was used to determine source significance assuming power-law spectra on top of standard diffuse model**
- **Define $TS = 2 \Delta \log(\text{likelihood})$ comparing models with and without the source. Cut at $TS = 25$, corresponding to about 4σ or $2.5E-5$ probability (4 degrees of freedom including source position).**



Works well at high latitudes.
Peak in TS distribution at threshold.
TS distribution flatter close to Galactic plane. Faint sources are not detected.

Sensitivity map

- Structure is mostly that of the interstellar medium
- Below 10^{-8} ph/cm²/s outside the Galaxy ($|b| > 30^\circ$)
- Strong dependence on spectral index

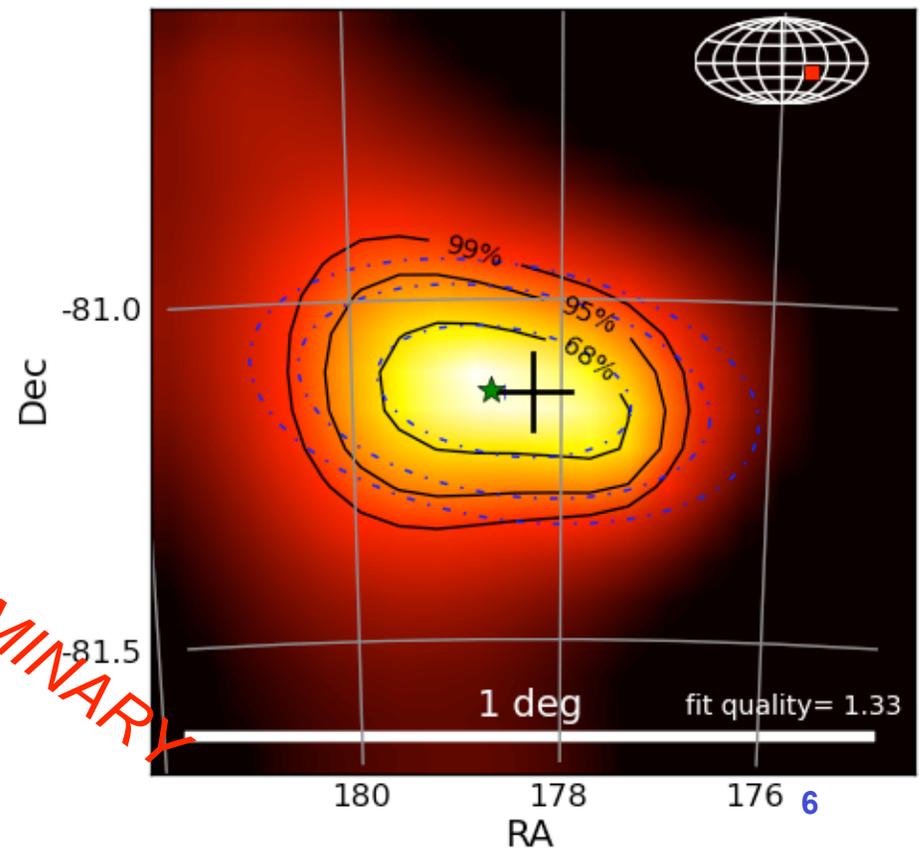
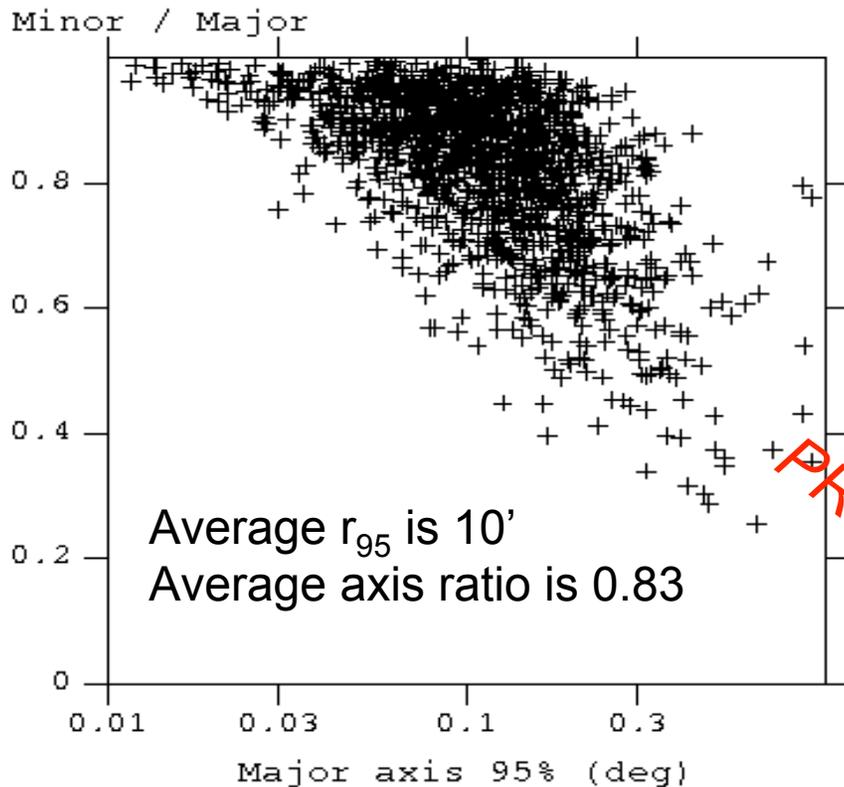


Flux > 100 MeV required to reach TS=25 for average $E^{-2.2}$ spectrum

Galactic coordinates, Aitoff projection

Source localization

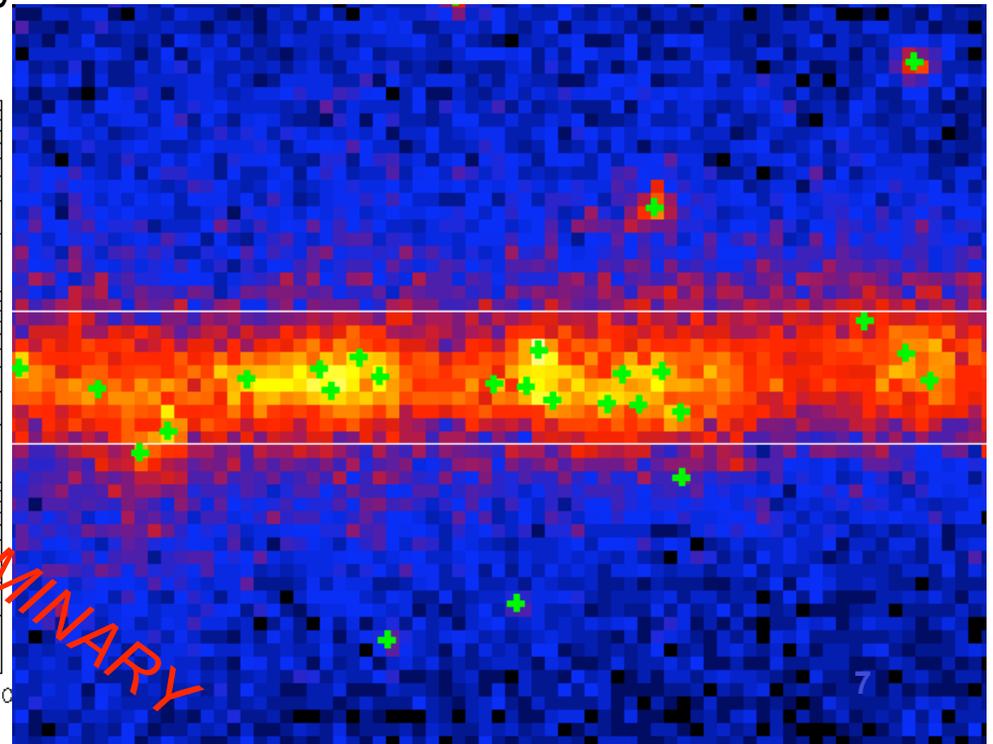
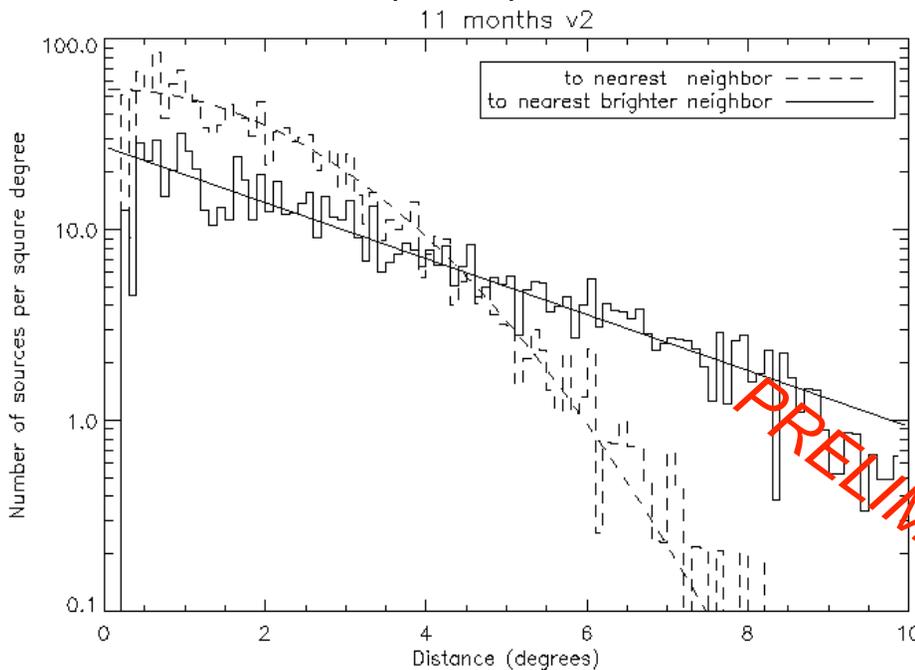
- Conservative error radii adjusted on known associations
- Conservative 0.012° absolute limit based on bright pulsars
- Elliptical parameters whenever can be extracted, mostly round.
- A difficult example is below. Cross is 1σ 1D error from other localization method for comparison



PRELIMINARY

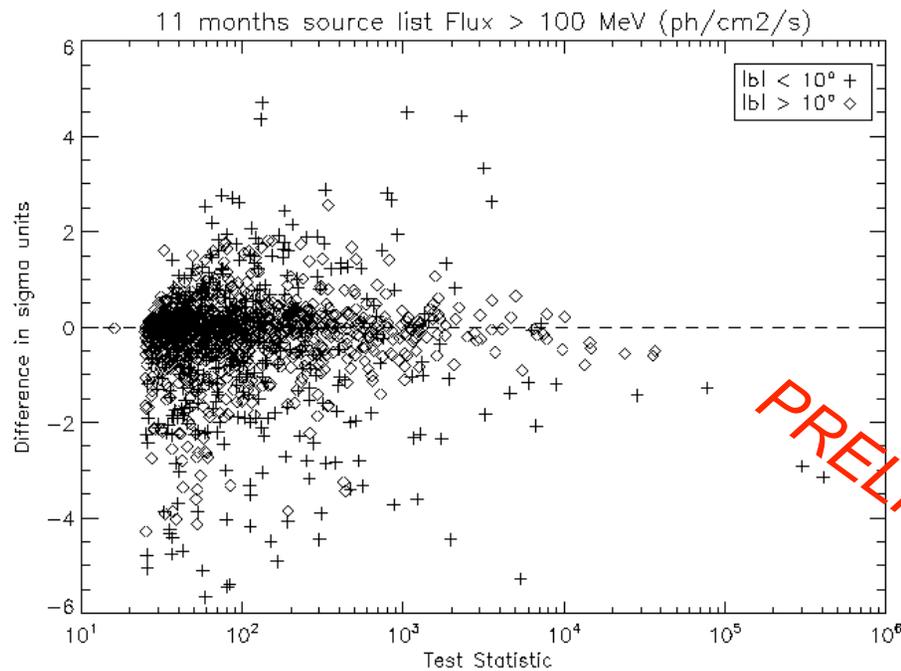
Source confusion

- Average distance between sources outside the plane is about 3°
- Much more than r_{68} at typical detection energy (0.8° at 1 GeV)
- Extrapolation to 0 distance indicates that only 6 sources or so are confused outside the plane
- Introduces additional bias against very soft sources ($\Gamma \geq 3$)
- **15° region of the Galactic ridge above 1 GeV**
- **Crosses are sources, pixel is 0.2°**
- **Sources not clearly separated**
- **Could be unmodeled diffuse emission**

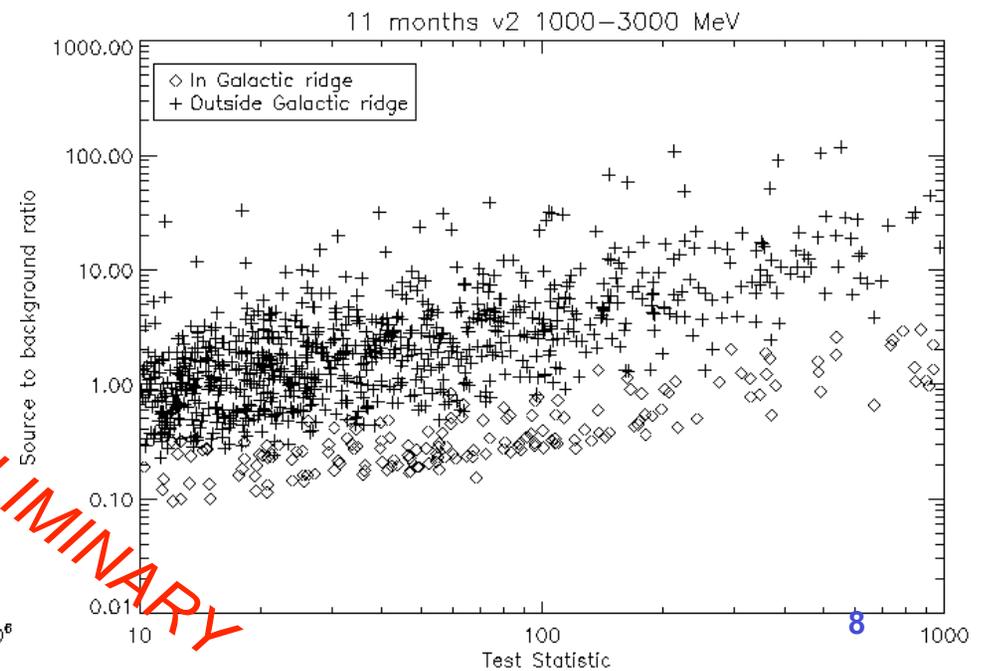


Diffuse emission uncertainties

- Use two different diffuse models to assess sensitivity of sources
- With good statistics, 10% of the background can be significant
- Inside Galactic plane, dispersion due to diffuse model is 1.8 sigma
- Outside plane, dispersion is 0.7 sigma
- **Source to background ratio within r_{68} is not very large in the Galactic plane even above 1 GeV.**
- **Has to go above 3 GeV to be above 50%**

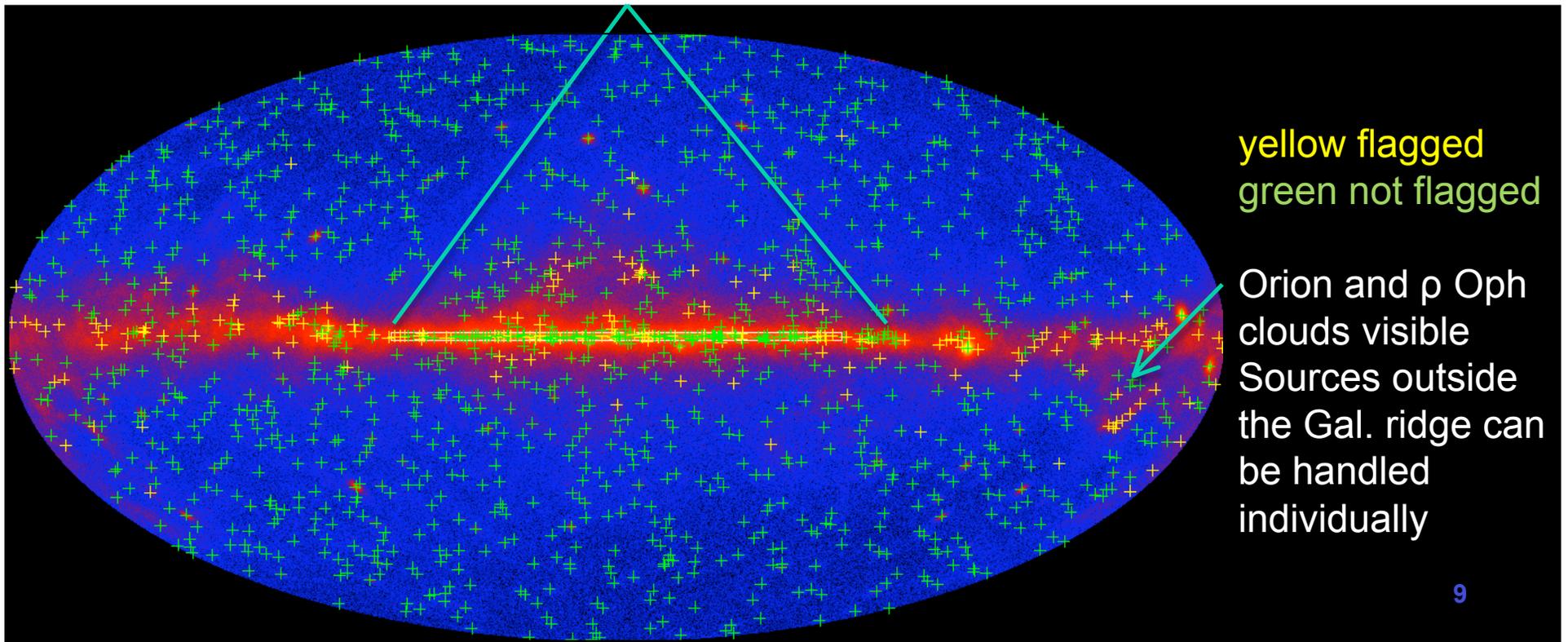


PRELIMINARY



First LAT source catalog: 1451 sources

- The Galactic ridge ($|\text{lat}| < 1^\circ$, $|\text{lon}| < 60^\circ$) has serious difficulties: sources are close to each other, are not high above the background below 3 GeV, and the Galactic diffuse model is very uncertain there. This even affects sources statistically very significant ($\text{TS} > 100$).
- We plan to flag Galactic ridge sources to warn against using them without detailed analysis. Of course there are still many true sources in there, including pulsars and SNRs.



First LAT Source Catalog Conclusions

- Typical 95% error radius is **10 arcmin**
- About 250 sources show evidence of **variability**
- About half the sources are **associated** positionally, mostly with blazars and pulsars
- **Other classes** of sources exist in small numbers (XRB, PWN, SNR, starbursts, globular clusters, radio galaxies, narrow-line Seyferts)
- Uncertainties due to the diffuse model, particularly in the **Galactic ridge**, should be kept in mind for low-latitude and local cloud studies
- The Catalog is an **analysis product** but also a **useful input** for many other studies with LAT data

On using the Catalog for Analysis

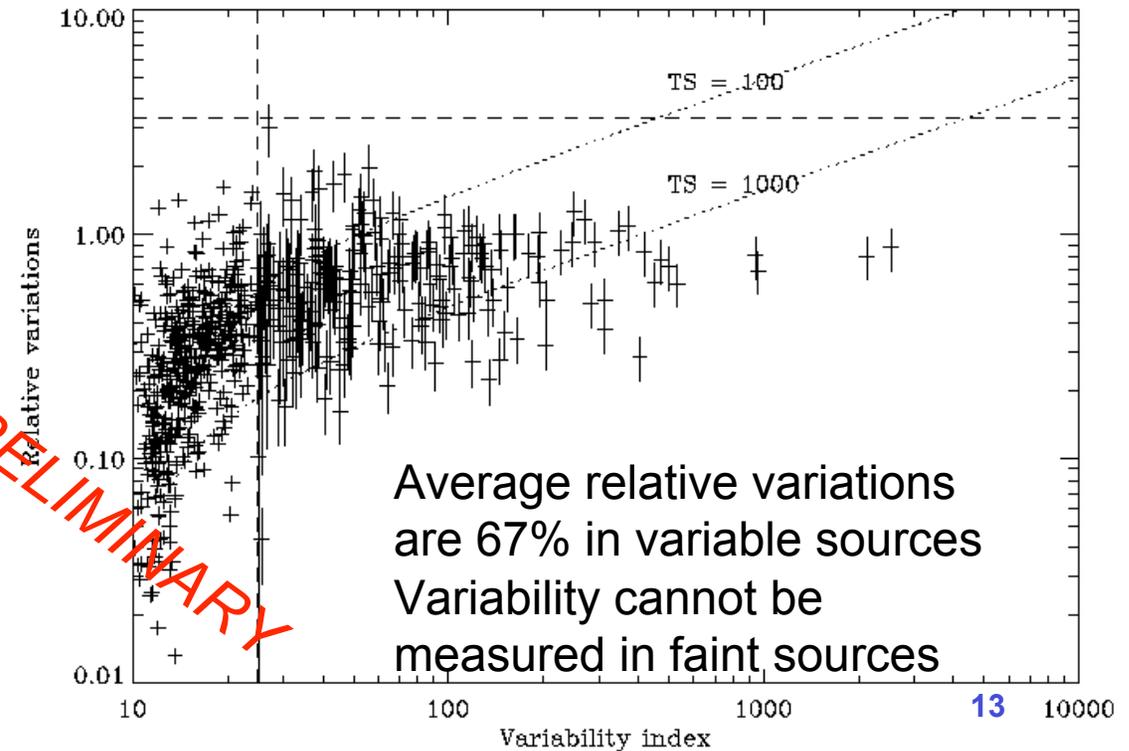
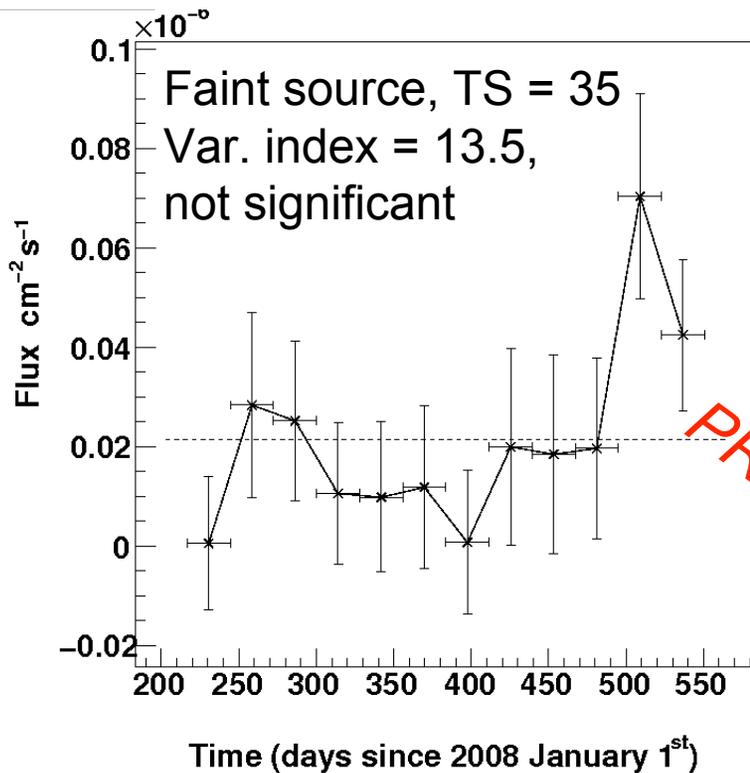
- **The 1FGL Catalog will be distributed via the FSSC as:**
 - **FITS table**
 - **XML input file for gtlite/ModelEditor**
 - **ds9 region file**
- **Useful for selecting specific sources for further study**
 - **More or less data**
 - **Different spectral models**
 - **Different diffuse emission models**
- **Useful for defining an initial point-source model for any ROI**
- **Additional items to note:**
 - 1) **The gamma-ray sky (blazars in particular) is highly variable. Flux and spectral information in the 1FGL are only time averages for the 11 months of the data set.**
 - 2) **For studies of diffuse emission at low latitudes and toward prominent local clouds, be aware that some 1FGL sources may be unresolved diffuse emission – whether they should be in the source model is a case-by-case decision.**



Backup Slides

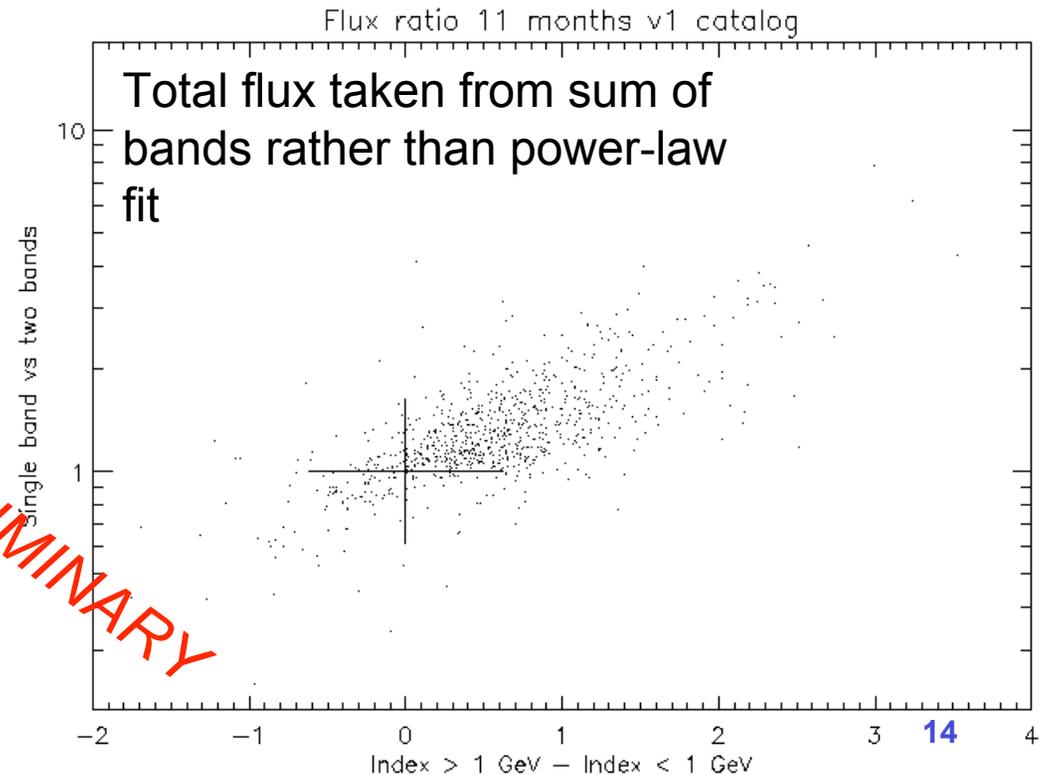
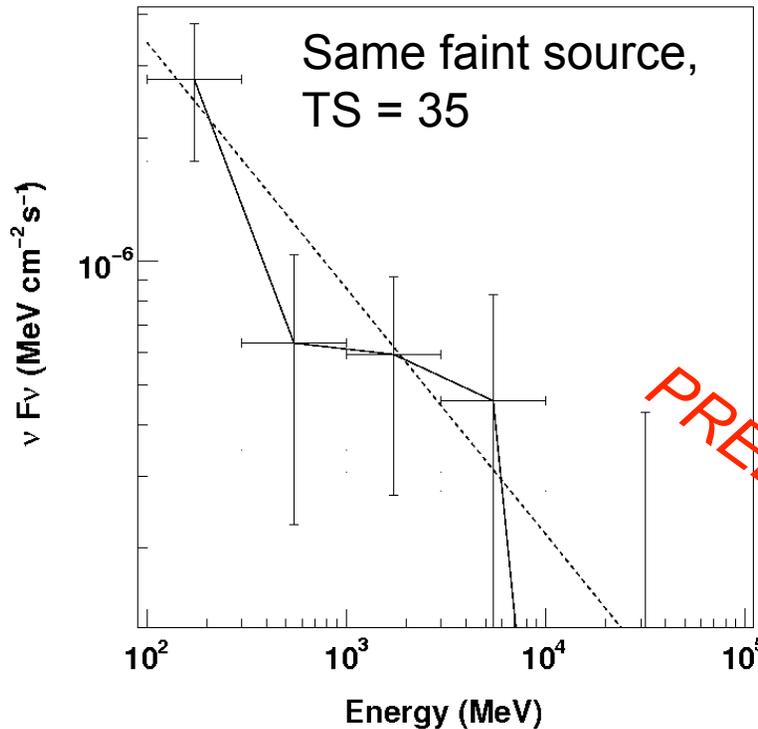
Source variability

- Build light curves of all sources on one-month time scale
- Pulsars are stable within 3%
- Bright blazars are very clearly variable
- Variability index: χ^2 against constant hypothesis. 250 variable sources
- Relative variations: $\Delta F/F$ where $\Delta F^2 = \text{measured variance minus Poisson variance}$



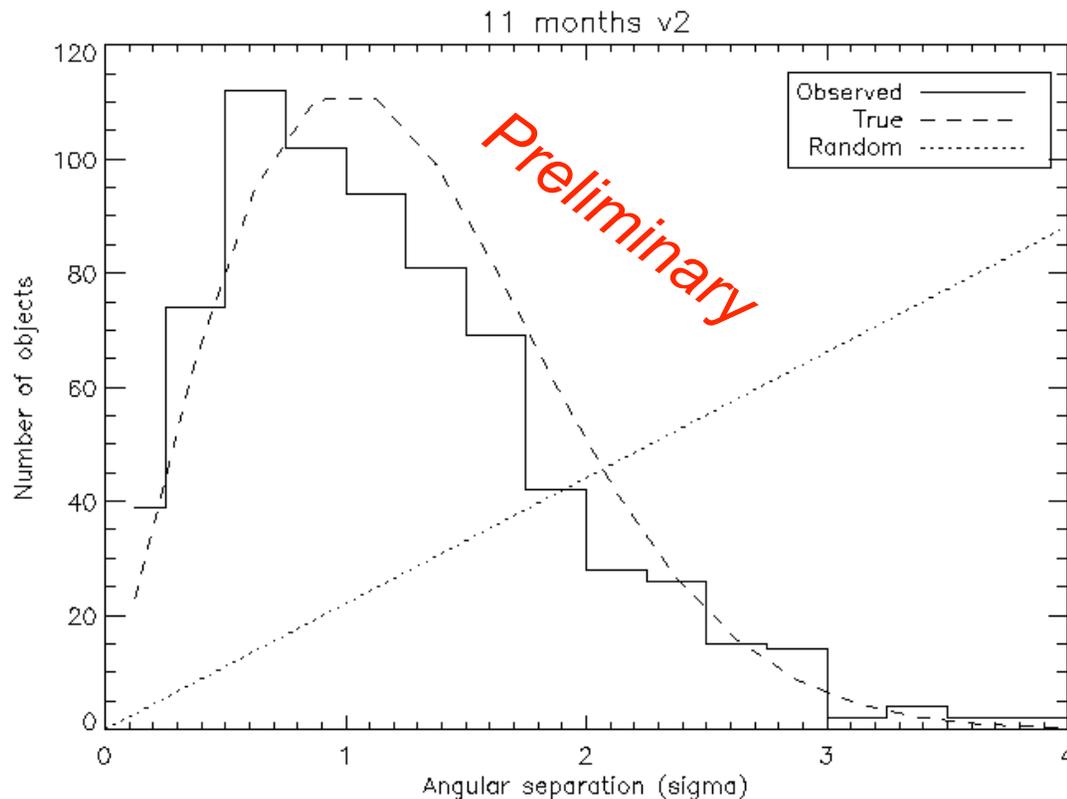
Source spectra

- Extract flux of all sources in 5 bands from 100 MeV to 100 GeV
- Upper limit in bands in which sources are not significant
- Average spectrum is broken, power-law estimate is too high
- Sources not significant in all bands, total flux not well measured (strong detection does not imply well measured $F_{>100 \text{ MeV}}$)



Source association

- Likelihood ratio between true association (gaussian distribution with width defined from r_{95}) and random association (flat at counterpart density)
- Typically one half of the sources are associated to a plausible counterpart (radio blazar, pulsar, PWN, SNR, XRB), down from 2/3 at 0FGL (brighter)



Point sources only
No doubt that most of these associations are true.
 r_{95} was multiplied by 1.2 to cover the tail.
The distance distribution may be more complex than a simple gaussian

Other LAT catalogs

- **The AGN catalog, drawn from sources above 10° latitude, excluding known non AGN. Goes with the full source catalog, which adopts the AGN associations. Similar to LBAS (ApJ 700, 597) vs BSL (ApJS 183, 46)**
- **The pulsar catalog, drawn from all pulsed detections of both radio and γ -ray pulsars. First version (6 months of data) submitted to ApJ (arXiv:0910.1608)**
- **The gamma-ray burst catalog, drawn from all GRB detections. Completely separate (bright GRBs are actually excluded for the LAT source catalog). Available in quasi real time at Fermi SSC**